

Example: Image Restoration



camera



- Imaging system introduces distortion
 - out of focus lens
 - atmospheric distortion
 - Hubble telescope mirrors
 - Motion blur
- Image restoration aims to reduce the distortion by post-processing the image

Model: approximate distortion with a linear shift-invariant system PSF $h[m,n]$

2

$$g[m,n] = h[m,n] * f[m,n]$$

observed image "blur" PSF "clean" image

↕ DSFT

$$G(u,v) = H(u,v) F(u,v)$$

Given $H(u,v)$ —

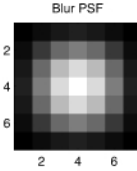
$$F(u,v) = \frac{G(u,v)}{H(u,v)} \quad \xleftrightarrow{\text{DSFT}} \quad f[m,n]$$

Cautions - watch for $H(u,v) = 0$

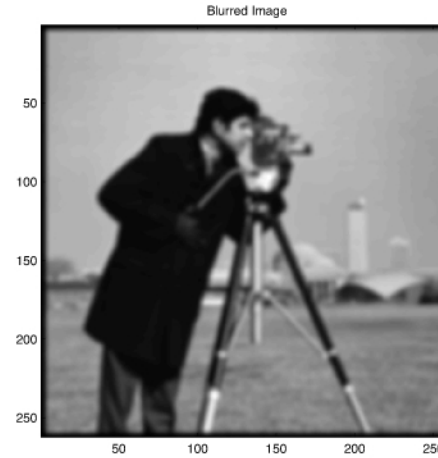
Model for Blur



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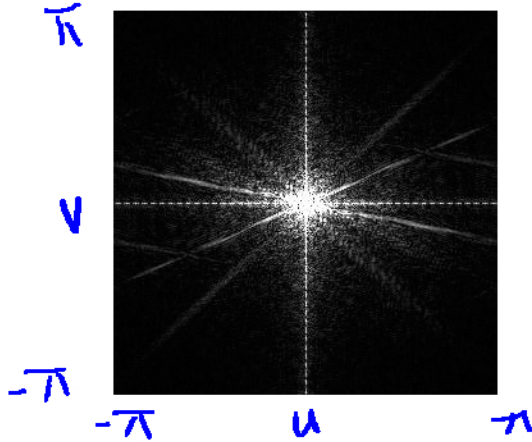
=



observed image

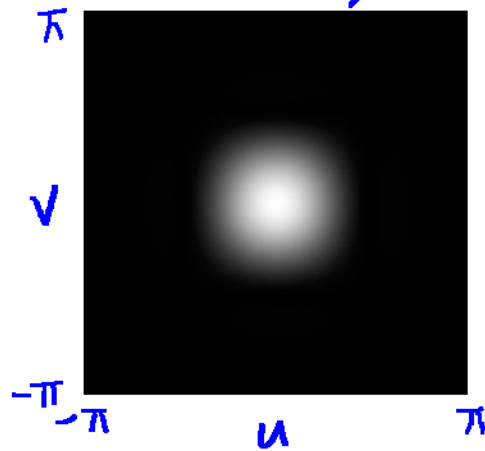
DSFT

$F(u,v)$



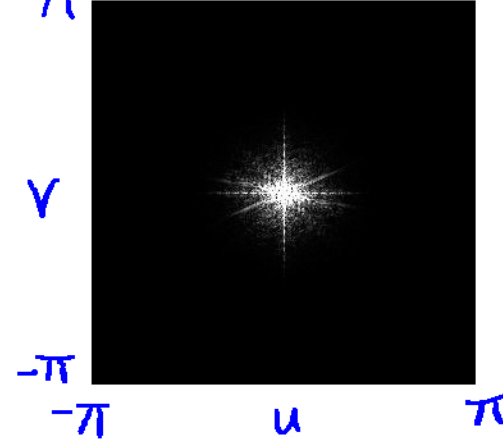
X

$H(u,v)$

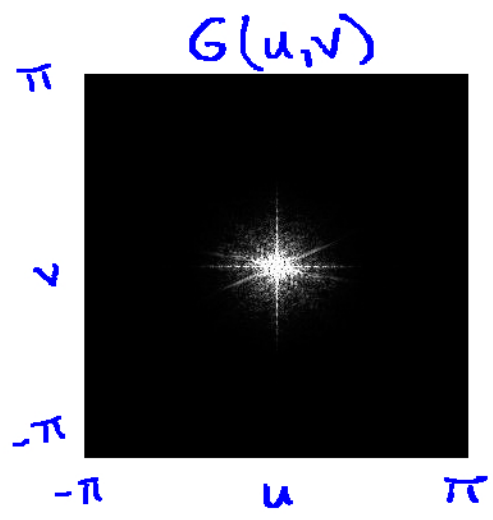


=

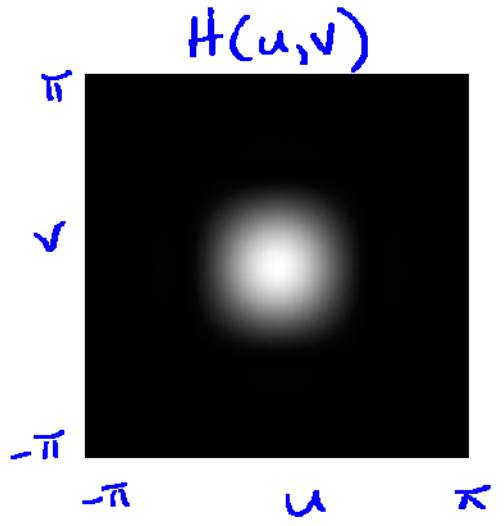
$G(u,v)$



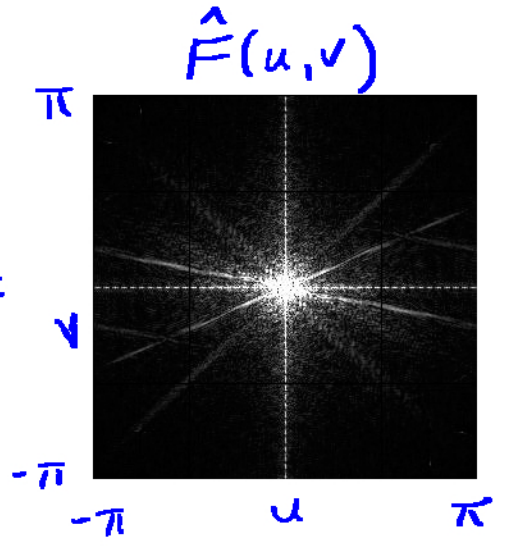
Deblurring



- /



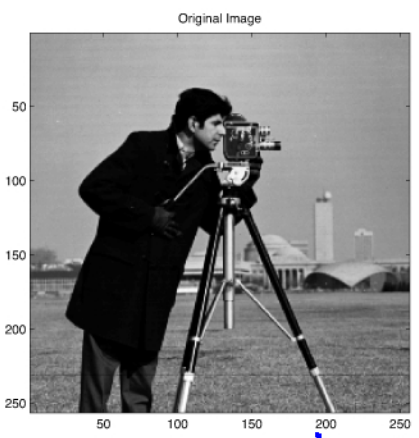
=



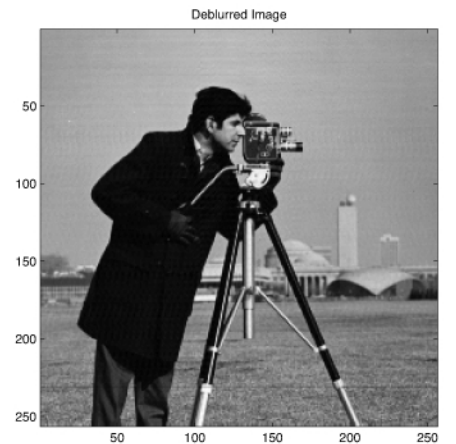
DSFT

Cautions

- must know $h[m,n]$
- zero or small $H(u,v)$
- Observation noise



original



deblurred

Observation Noise

$$g[m,n] = h[m,n] * f[m,n] + w[m,n]$$

↙ noise

$$\hat{F}(u,v) = \frac{G(u,v)}{H(u,v)} = \frac{H(u,v)F(u,v)}{H(u,v)} + \frac{W(u,v)}{H(u,v)}$$

small values of $H(u,v) \rightarrow$ amplify noise

limit $\frac{1}{H(u,v)}$

